

The background of the slide is a photograph of a large radio telescope array in a flat, open landscape. The foreground and middle ground are filled with numerous white, conical antenna structures. In the distance, a dark-colored pickup truck is parked on a dirt road. The sky is a clear, pale blue with a few wispy clouds. The overall scene is brightly lit, suggesting a sunny day.

# Embedded Element Patterns in Hierarchical Calibration of Large Distributed Arrays

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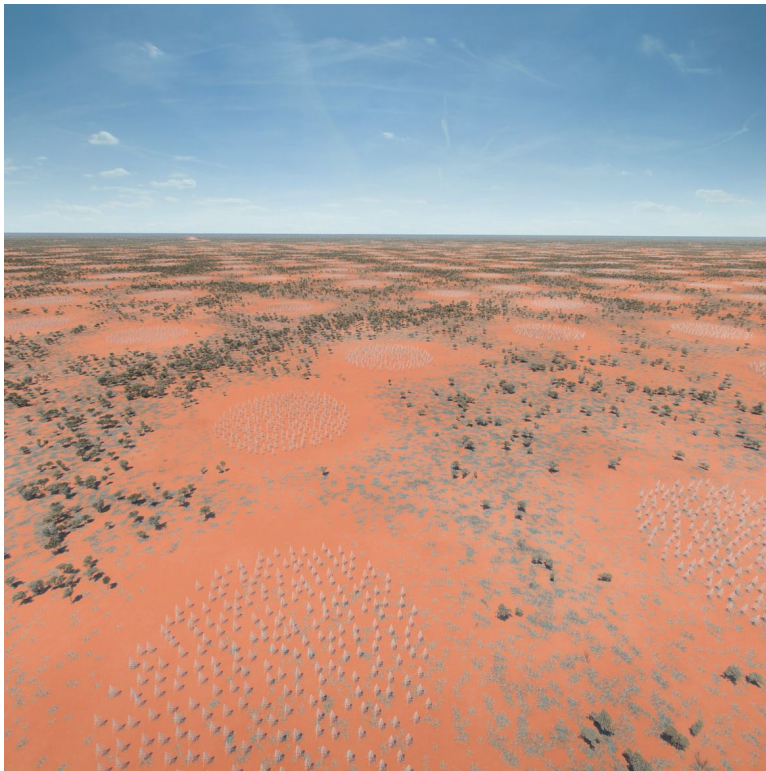
URSI GASS 2020

# SKA-LOW and LOFAR

**Low**-frequency instrument of  
**S**quare **K**ilometre **A**rray

50 – 350 MHz

**Array of subarrays (stations)**



**Low Frequency Array**

10 – 250 MHz

**Array of subarrays (stations)**



# Hierarchical calibration

## Calibration at station level needed for

- Accurate beamforming
  - Ensures station sensitivity
  - Allows beam shaping, e.g., nulling



## Calibration at array level needed for

- High-dynamic range imaging
  - Needs station beam stability
- Absolute calibration
  - Flux transfer from flux calibrators



# Embedded Element Patterns

Virone et al., IEEE TAP, 2018

Di Ninni et al., IJAP, 2019

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## EEPs can be simulated and validated in-situ using drones

### Simulation can compute (in order of increasing costs)

- Isolated EEP: EEP of isolated antenna
- Average EEP: average EEP of all antennas in station
- Individual EEPs: different EEP for each antenna within station

### Questions

- What is needed for station-level calibration?
- What is needed for array-level calibration?

### Derived question

- What calibration accuracy is needed?

Beamforming efficiency with RMS phase error  $\sigma_\varphi$

$$\eta_{\text{BF}} = \cos^2(\sigma_\varphi)$$

Implications:  $\eta_{\text{BF}} \geq \{0.99, 0.98\}$  requires  $\sigma_\varphi \leq \{5.7, 8.1\}$  degrees

Beamforming eff. with relative RMS error  $\varepsilon$  on real and imaginary part

$$\eta_{\text{BF}} = \frac{1}{1+2\varepsilon^2} \approx 1-2\varepsilon^2$$

Implications:  $\eta_{\text{BF}} \geq \{0.99, 0.98\}$  requires  $\varepsilon \leq \{0.071, 0.10\}$

Array level calibration needs to be able to track dir. dep. gain changes

First order model for varying gain of  $i$ th station  $g_i = g_{0,i} + \alpha_i t$

To keep errors below 20% of thermal noise, we need

$$\frac{|\alpha_i|}{|g_{0,i}|} \leq \sqrt{\frac{12}{5}} \frac{1}{\sqrt{\text{SNR}}} \frac{1}{\tau}$$

where  $\tau$  is the calibration interval in which the given SNR is achieved

Example: SNR = 10 and  $\tau = 600$  s allows rate of change of 0.082%/s

Note: time needed to achieve a certain SNR depends on SEFD

**Hence: more sensitive instrument (lower SEFD) can keep up with faster gain changes**

# Impact of ignoring EEPs (1)

Wijnholds, SKA-LOW meeting, Florence, 2019

Haslam et al., A&A Suppl, 1982

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Simulation setup for 256-element SKA-LOW station

- Mock data based on simulated EEPs and Haslam map
- Calibration model assuming identical EEPs equal to average EEP
- Nominal gain equal to unity for each element
- 200 scenarios spread over 24 hours (one solution per 7.2 min)
- Simulation done for both SKALA4AL and EDA at 110 MHz
- Gain solutions used to calculate AF for each instant

$$AF(\mathbf{l}; \mathbf{l}_0) = \mathbf{w}^H(\mathbf{l}_0)(\mathbf{g} \odot \mathbf{a}(\mathbf{l})) = \sum_{p=1}^P g_p \exp\left(\frac{-2\pi i}{\lambda} \mathbf{x}_p(\mathbf{l} - \mathbf{l}_0)\right)$$

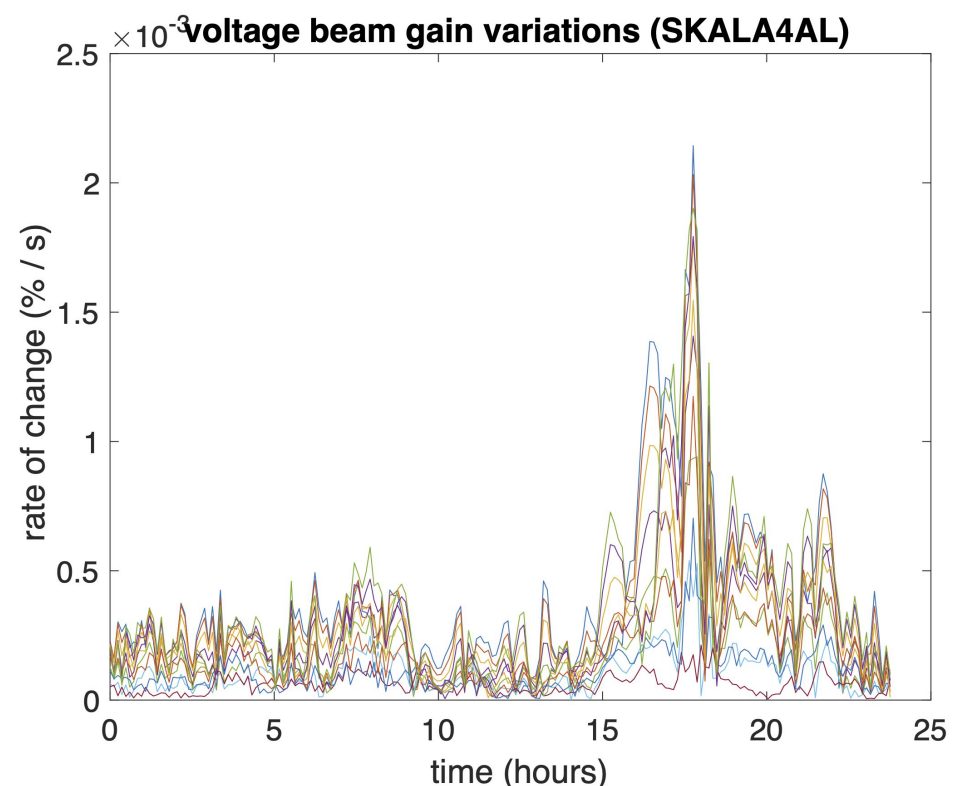
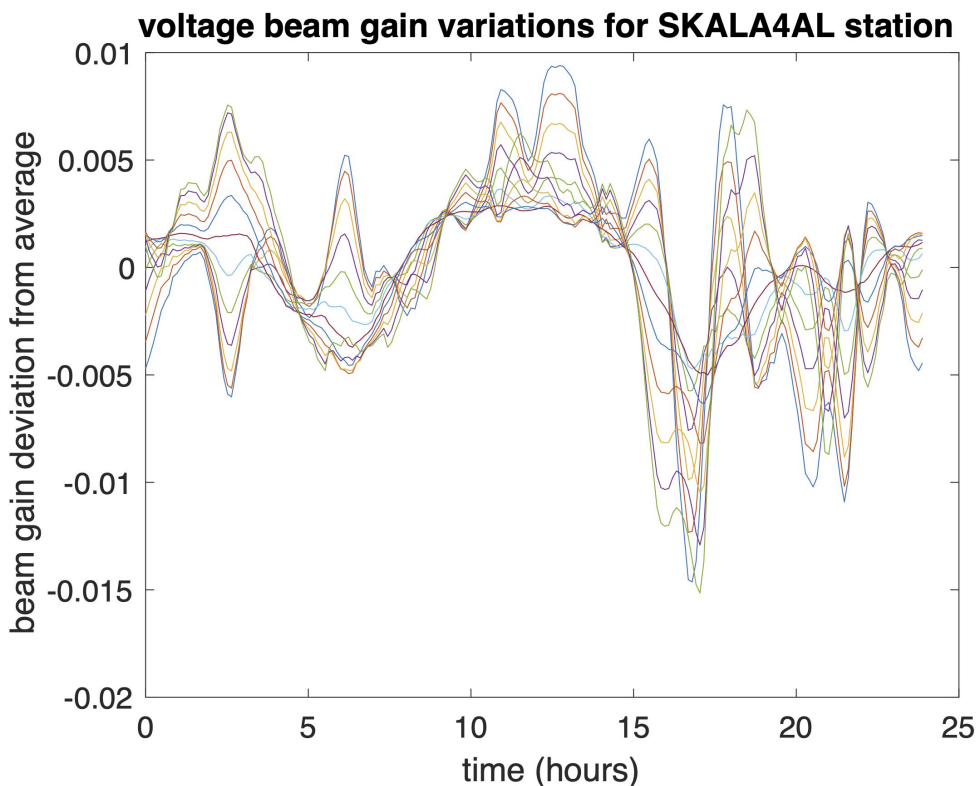
- Average AF normalized to have unit peak gain

# Impact of ignoring EEPs (2)

Left: beam gain variations along cross-section through station main beam with largest variations

Right: rate of change at each point of this cross-section

**Conclusion: average EEP sufficient if sky model is correct**





Balancing against absolute flux calibrator accuracy:

- Typical absolute flux accuracy of flux calibrators is  $\sim 5\%$
- Instrument should not be limiting, so LOFAR2.0 has set reproducibility of absolute flux calibration at 2%
- Here, reproducibility applies to the absolute flux calibration in the target field for different calibrators or the same calibrator at different sidereal times
- Assuming error towards calibrator and target field are uncorrelated gives tolerance of 1.4%

# Compliance assessment flux transfer

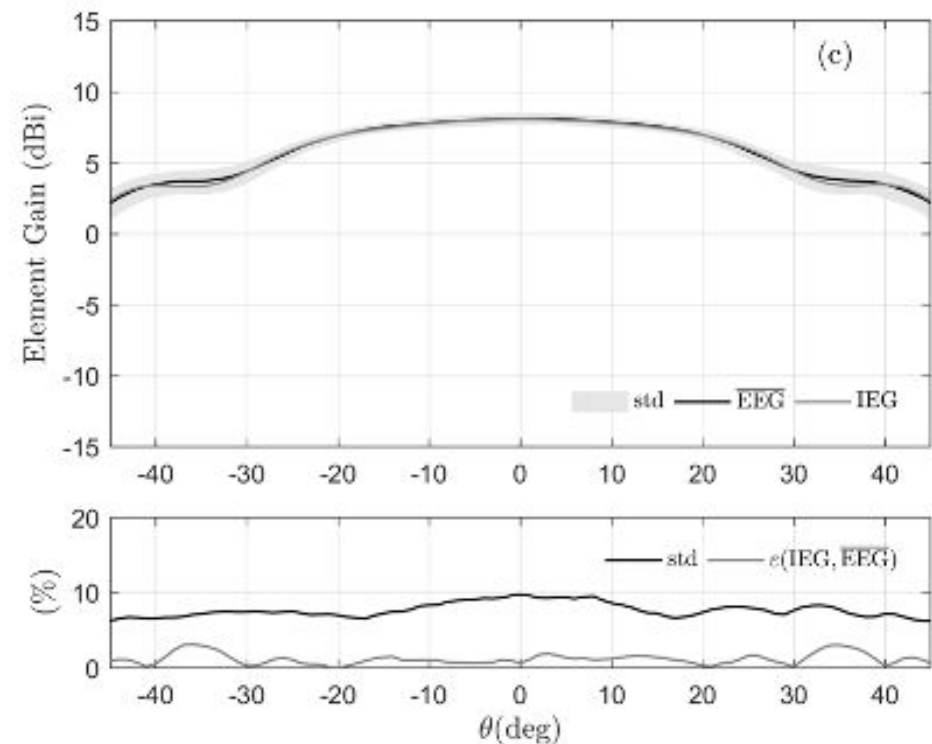
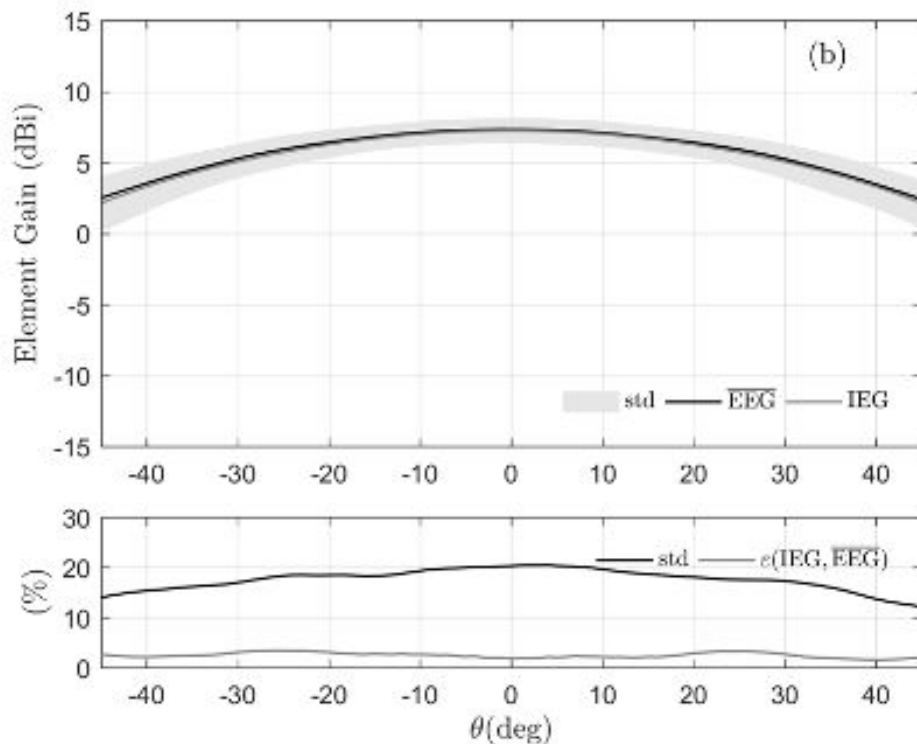
Di Ninni et al., EuCAP 2019

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Comparison between average EEP and isolated EEP for SKA-LOW

Patterns (top) and difference (bottom) at 110 (l) and 350 (r) MHz

Differences up to about 3%, average EEP needed to meet requirement



## Station level

- Requirement proposed on coherence during beamforming
- Requirement proposed on tolerable rate of change
- Both requirements can likely be met with an average EEP

## Array level

- Requirement proposed on reproducibility of absolute flux calibration
- SKA-LOW needs average EEP to satisfy this requirement

**Individual EEPs may (fortunately) not be necessary**